

1 The Unique, Well Posed Reduced System for  
2 Atmospheric Flows: Robustness In The Presence  
3 Of Small Scale Surface Irregularities

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5 Dedicated to my mentor and colleague, Heinz-Otto Kreiss

6 **Abstract**

7 Numerical analysis requires that a number of derivatives of the contin-  
8 uum solution of any differential system of equations exist in order that the  
9 numerical approximations of the derivatives of that system are ensured to  
10 have sufficiently small truncation errors. If a hyperbolic partial differen-  
11 tial system of equations also contains multiple time scales (as is the case  
12 for the atmospheric equations of motion) and it is the goal to accurately  
13 compute the component of the solution that evolves on the time scale  
14 of the advective terms (the component with the majority of the energy),  
15 then additional restrictions are required of the derivatives. For the initial  
16 value problem, deriving higher order time derivatives of the continuum  
17 solution using the Bounded Derivative Theory (Kreiss 1979 , Kreiss 1980)  
18 will lead to spatial elliptic constraints that must be satisfied to ensure  
19 the ensuing solution will evolve on a space and time scale of order unity  
20 in the scaled system. For the initial/boundary problem, the boundary  
21 conditions for the elliptic constraints must be derived from the well posed  
22 boundary conditions of the original hyperbolic system that ensure that

23 the ensuing solution in the limited area evolves on space and time scales of  
24 order unity. If these requirements are met, then the  $L_2$  energy estimates  
25 of the solution and a number of its spatial and temporal derivatives are  
26 independent of the fast time scales and ensure that the resulting limit  
27 system (the reduced system) as the fast scales approach infinity will auto-  
28 matically be well posed. In this manuscript the reduced system for large  
29 scale atmospheric flows will be introduced and a special property of the  
30 corresponding elliptic equation for the vertical component of the velocity  
31 will be discussed. In particular, the solution of that elliptic equation is  
32 not sensitive to small scale perturbations at the lower boundary so it can  
33 be used all of the way to the surface.

## 34 1 Introduction

35 The five time dependent partial differential equations for entropy(1), mass(1)  
36 and momentum(3) describe the evolution of many different kinds of fluids. The  
37 natural mathematical question is how does one understand the behavior of a  
38 particular kind of fluid exhibiting a particular kind of behavior. This question  
39 has been answered by introducing a scale analysis of the particular kind of  
40 flow of the fluid, i.e., by introducing characteristic variables that describe the  
41 typical values of the independent and dependent variables of that flow and then  
42 making a simple change of variables using those values. This technique is helpful  
43 in identifying the relative sizes of the individual terms in any given equation  
44 and has been used in many scientific areas including meteorology (Charney  
45 1948), oceanography (Browning, Holland, and Worley 1989) and plasma physics  
46 (Browning and Holzer 1992). However, that scale analysis does not ensure that  
47 the given flow will continue to evolve in time with the same chosen scales of  
48 motion.

49 Kreiss (Kreiss 1979 , Kreiss 1980) introduced the Bounded Derivative Theory

50 (BDT) for scaled (nondimensional) hyperbolic systems of equations with the  
51 advective terms of order unity (space and time scales of order of the advective  
52 terms) and off diagonal terms much greater than unity contributing to high  
53 frequency (fast time scale) components of the solution. To ensure that the  
54 ensuing solution would evolve on the order of the advective terms, i.e., of order  
55 unity, the mathematical  $L_2$  energy method applied to the solution and its space  
56 and time derivatives must yield norms of those functions on the order of unity  
57 for a time period on the order of the time scale of the advective terms. As these  
58 estimates of the solution and its space and time derivatives are independent  
59 of the fast (high frequency) time scales, the estimates hold as the size of the  
60 large terms increase to infinity. Thus the system that represents this limit, the  
61 reduced system, also satisfies these so estimates is automatically well posed and  
62 accurately describes the motion of interest.

63 A review of the BDT for the atmospheric equations for large scale flows in  
64 the midlatitudes is presented in Section 2. The scaled system from Browning  
65 and Kreiss (Browning and Kreiss 1986) is reproduced to reveal how the scaling  
66 produces a nondimensional system with large off diagonal terms. A simple  
67 example is used to show how the space and time derivatives become coupled  
68 so that both must be estimated to ensure a slowly evolving solution. Section  
69 3 introduces the reduced system for large scale midlatitude flows, although it  
70 has been shown that the reduced system also accurately describes mesoscale  
71 flows (Browning and Kreiss 2002). Numerical examples are presented in Section  
72 4. The examples use a heating function that is large scale in space and time  
73 to generate an evolving large scale solution. The solutions from the model  
74 based on the multiscale system and the model based on the reduced system are  
75 compared. In contrast to Richardson's equation for the vertical velocity in the  
76 primitive (hydrostatic) equations, it is demonstrated that the solution of the

77 elliptic equation for the vertical velocity in the reduced system is not sensitive  
78 to small scale noise at the lower boundary.

## 79 **2 Bounded Derivative Theory Review**

To determine the relative size of individual terms in a given equation of the partial differential system that describes large scale atmospheric motions, a simple change of variables is used. The characteristic scales of the independent and dependent variables describing the motion are used for this purpose, e.g., a horizontal length scale  $L = 1000$  km, a depth scale of  $D = 10$  km, a time scale  $T = 86400$  sec (1 day), a horizontal velocity scale  $U = 10$  m/s and a vertical velocity scale  $W = .01$  m/s. The pressure and density are scaled as perturbations of a mean state in hydrostatic equilibrium. For large scale motion in the atmosphere this leads to the following scaled (nondimensional) system of equations (Browning and Kreiss 1986):

$$\frac{ds}{dt} - \tilde{s}(w - H) = 0, \quad (2.1a)$$

$$\frac{du}{dt} + \epsilon^{-1}(\rho_0^{-1}p_x - fv) = 0, \quad (2.1b)$$

$$\frac{dv}{dt} + \epsilon^{-1}(\rho_0^{-1}p_y + fu) = 0, \quad (2.1c)$$

$$\frac{dw}{dt} + \alpha\epsilon^{-6}(\rho_0^{-1}p_z + \tilde{p}p + gs) = 0, \quad (2.1d)$$

$$\frac{dp}{dt} + \epsilon^{-1}wp_{0z} + \epsilon^{-2}\gamma p_0(u_x + v_y + \epsilon w_z) = 0, \quad (2.1e)$$

80 where  $d/dt = \partial/\partial t + u\partial/\partial x + v\partial/\partial y + \epsilon w\partial/\partial z$ . The nondimensional depen-  
81 dent variables  $s, u, v, w, p$  are the reciprocal of entropy (hereafter referred to as  
82 entropy for brevity), velocity components, and pressure perturbation from the  
83 mean, respectively. The dimensionless functions  $\rho_0(z)$  and  $p_0(z)$  are the mean  
84 hydrostatic state values of the density and pressure,  $s_0 = \rho_0 p_0^{-1/\gamma}$ ,  $\tilde{s} = 10s_{0z}/s_0$ ,

85  $\tilde{p} = -p_{0z}/(\gamma\rho_0p_0)$  and  $f(y)$  is the Coriolis term. The nondimensional con-  
 86 stant  $g$  is the gravitational constant and  $\gamma = 1.4$ . The dimensionless function  
 87  $H(x, y, z, t)$  is essentially the sum of all heating and cooling sources. Inverse  
 88 powers of  $\epsilon = 1/10$  represent large terms. (Note that the original dimensional  
 89 equations can be obtained by setting  $\epsilon = 1$ .) For the original scaling  $\alpha = 1$ ,  
 90 but for the multiscale system it is  $\alpha = (D/L)^2$  which is  $\epsilon^4 = 10^{-4}$ . The latter  
 91 value of  $\alpha$  has been proved mathematically to reproduce the slowly evolving in  
 92 time solution of (2.1) with  $\alpha = 1$  to at least two digits of accuracy. Because  
 93 there are five time dependent equations there are five different frequencies: one  
 94 associated with advective motions, two with inertial/gravity waves and two with  
 95 sound waves (Browning and Kreiss 1985). The BDT theory was developed to be  
 96 able to select the five initial conditions so that the components of the solution  
 97 that are associated with the latter four frequencies would remain small for a  
 98 given period of time.

99 The first such scaling was performed by Charney without the mean hy-  
 100 drostatic state removed (Charney 1947, Charney 1948) and subsequently by  
 101 Browning and Kreiss (Browning and Kreiss 1986) with the mean hydrostatic  
 102 state removed. Charney discovered that there were two large terms that were  
 103 8 orders of magnitude larger (6 orders with the mean state removed) than the  
 104 remaining terms in the time dependent equation for the vertical component of  
 105 velocity,  $dw/dt$ , that were impossible to compute accurately using numerical  
 106 methods of the time. Instead of doing so, the two terms, the vertical pressure  
 107 gradient and the gravitational term, were set to be equal. This equality between  
 108 these two terms is called hydrostatic balance and leads to a modification of the  
 109 system called the primitive equations. The resulting columnar integral equation  
 110 at each horizontal point for the vertical velocity is called Richardson's equation.

111 A scale analysis does not by itself prove that the motion will evolve as de-

112 scribed by the characteristic scales of motion. This must be done by mathemat-  
 113 ics using the theory of hyperbolic systems of equations in conjunction with the  
 114 Bounded Derivative Theory (Kreiss 1979, Kreiss 1980). To determine the subse-  
 115 quent motion of such a system requires estimates of the ensuing in time spatial  
 116 and temporal derivatives. The estimates for these derivatives are determined by  
 117 differentiating the equations with respect to space or time as appropriate and  
 118 then using the mathematical  $L_2$  energy method to estimate the norms of the  
 119 derivatives at a later time to ensure that the solution will continue to evolve on  
 120 the slow time scale. An important detail in such arguments can be considered  
 121 using the so called Kreiss equation

$$u_t = a(x, t)u_x. \quad (2.2)$$

122 To estimate the ensuing time derivative of a solution of this equation differen-  
 123 tiate the equation with respect to time

$$(u_t)_t = a(x, t)(u_t)_x + a_t(x, t)u_x. \quad (2.3)$$

124 Note that  $u_t$  satisfies the same equation as  $u$  with the exception of one term.  
 125 That term couples the space and time derivative terms. Thus in the BDT the  
 126 energy estimates must show that both the higher order spatial and temporal  
 127 derivatives evolve as specified in the scaling, namely, with the space scale  $L$   
 128  $= 1000$  km and the time scale on the order of  $T = 1$  day, or the solution  
 129 will not continue to evolve in the chosen manner. This requirement precludes  
 130 the primitive equation solution from evolving correctly because the columnar  
 131 equation for the vertical component of the velocity can change discontinuously  
 132 from horizontal point to point because of switches in the heating parameteri-  
 133 zations. Those discontinuities violate the spatial derivative estimates required

134 by the BDT. They also require unrealistically large dissipation because they  
 135 inject energy into the smallest scales in a numerical model. That dissipation  
 136 reduces the numerical accuracy of a numerical method by orders of magnitude  
 137 (Browning, Hack, and Swarztrauber 1989, Browning and Kreiss 1994 ). Note  
 138 that the initial-boundary value problem for the primitive equations is not well  
 139 posed (Oliger and Sundström 1978) and that also indicates that it is not the  
 140 correct reduced system.

141 Using the BDT to initialize the unmodified hyperbolic Euler equations with  
 142 the appropriate space and time derivatives ensures that the evolution of the  
 143 solution on the chosen scales will require no dissipation and minimal numerical  
 144 accuracy as will be shown in the numerical examples to follow. As has been dis-  
 145 cussed before (Browning and Kreiss 2002), the elliptic initialization constraints  
 146 can be used in conjunction with a time dependent equation for the vertical com-  
 147 ponent of vorticity to form an automatically well posed system that accurately  
 148 describes the evolution of the large scale motion.

### 149 **3 Reduced System**

150 In this section the reduced system for large scale atmospheric motions will be  
 151 described. As mentioned previously, a time dependent equation for the vertical  
 152 component of vorticity (the only variable that can be used globally for the time  
 153 dependent slowly evolving variable in time and space for all scales of motion)  
 154 is added to the initialization constraints for the multiscale system. The time  
 155 dependent equation for the vertical component of vorticity  $\zeta = -u_y + v_x$  can  
 156 be derived by appropriately cross differentiating equations (2.1a) and (2.1b):

$$\frac{d\zeta}{dt} + v_z w_x - u_z w_y + (f + \delta)\zeta + f_y v = 0, \quad (3.1)$$

157 where  $\delta = u_x + v_y$  is the horizontal component of divergence. The elliptic  
 158 initialization constraints for  $s$  and the vertical velocity  $w$  are (Browning and  
 159 Kreiss 2002)

$$\nabla^2 s = -\{\rho_0[f\zeta - f_y u + 2(u_x v_y + u_y v_x)]\}_z / (\rho_0 g), \quad (3.2)$$

$$\nabla^2 w + f^2 (g\tilde{s})^{-1} [w_{zz} + \rho_0 z (\rho_0)^{-1} w_z] = \nabla^2 H - (g\tilde{s})^{-1} R_1, \quad (3.3)$$

$$R_1 = -gC_2 - fC_1 \rho_0, \quad (3.4)$$

$$C_1 = u_z (\rho_0 \zeta)_x + v_z (\rho_0 \zeta)_y, \quad (3.5)$$

$$C_2 = u_{xx} s_x + 2u_x s_{xx} + u_{yy} s_x + 2u_y s_{xy} \quad (3.6)$$

$$+ v_{xx} s_y + 2v_x s_{xy} + v_{yy} s_y + 2v_y s_{yy}, \quad (3.7)$$

160 where  $C_1$  and  $C_2$  are the commutators derived previously (Browning and Kreiss  
 161 2002). The quantity between French braces in the equation for  $s$  is essentially  
 162 just the vertical derivative of the right hand side of the nonlinear balance equa-  
 163 tion, i.e., the equation is derived by using the two dominant terms of hydrostatic  
 164 balance from equation (2.1d). (The additional  $\tilde{p}$  term can be added in a similar  
 165 manner.) The horizontal smoothing of the right-hand-side of the balance equa-  
 166 tion is retained in the elliptic equation for  $s$ , but there is no smoothing of the  
 167 vertical derivative. However, in the equation for  $w$ , there is vertical smoothing  
 168 and that is what results in the well posedness of the reduced system. By using  
 169 the equation for  $s$  instead of the one for  $p$ , the mean of  $s$  is not required as only  
 170 derivatives of  $s$  appear in the right-hand-side for  $w$ . Note that we have neglected  
 171 a number of terms of order  $\epsilon$  in the derivation of the equation for  $w$  to simplify  
 172 the presentation. If required, they can be added by a simple iterative method.

173 The equation for  $w$  has several very special properties, namely, that small



174 scale perturbations of the lower boundary condition have only a minor impact on  
 175 the solution while larger scale perturbations do. This is physically important  
 176 in both cases, e.g., the so called lake effect on large scale storms. Thus the  
 177 equation for  $w$  can be used all of the way to the surface without the need for the  
 178 ad hoc discontinuous boundary layer parameterization to artificially slow down  
 179 the unrealistic growth of the velocity at the surface when using Richardson's  
 180 equation (Sylvie Gravel, personal communication). Note that the equation for  
 181  $w$  is similar to the quasi-geostrophic  $\omega$  equation (Charney 1947, Charney 1948).

182 The horizontal divergence in the reduced system is given by the balance  
 183 between the large terms in equation (2.1e)

$$\delta = -[w_z + wp_{0z}(\gamma p_0)^{-1}]. \quad (3.8)$$

184 Given the vorticity  $\zeta$  and the divergence  $\delta$ , the horizontal components of velocity  
 185 must be computed from the Helmholtz equations

$$\nabla^2 u = -\zeta_y + \delta_x, \quad (3.9)$$

$$\nabla^2 v = \zeta_x + \delta_y, \quad (3.10)$$

186 in order to connect these constraints to the well posed boundary conditions for  
 187 the hyperbolic system (2.1) that ensure a slowly evolving solution in a limited  
 188 area.

## 189 4 Numerical Examples

190 The details of the numerical approximation of the multiscale system in a channel  
 191 2000 km square and 12 km high have been presented earlier (Browning and

192 Kreiss 2002) so here we just summarize that method. The multiscale equations  
193 were approximated by the leapfrog method in space and time. The spatial  
194 derivatives needed no special treatment in  $x$  because the solution was periodic  
195 in that direction. At the north and south wall boundaries the  $y$  component of  
196 velocity  $v = 0$  so the boundaries were treated with inflow/outflow conditions  
197 with the  $y$  component of the velocity in that treatment set to 0. Similarly for  
198  $w$  at the bottom and top boundaries.

199 For the reduced system, the fourth order Runge-Kutta method in time and  
200 second order centered differences in space are used for the vorticity equation with  
201 the diagnostic quantities determined at each stage using the elliptic equations  
202 for  $(u, v)$ ,  $s$ ,  $w$  then  $(u, v)$  again in that order. No special treatment is needed  
203 for the vorticity equation at either the north or south boundaries. Note that in  
204 equation (2.1a) for  $s$ , at the boundaries at the bottom and top of the channel  
205  $w = 0$ . Thus, if the initial value of  $s$  is zero and there is no heating on those  
206 boundaries,  $s$  will remain identically zero there even as the horizontal velocities  
207 become nonzero. If there is initial horizontal velocity on those boundaries,  $s$   
208 will be horizontally advected with the heating (if any) acting as a forcing term.  
209 In either case this provides the variable  $s$  at the bottom and top of the channel.  
210 In the equation for  $u$  the boundary condition  $-u_y = \zeta$  is used at the north and  
211 south boundaries, while  $v$  is identically zero there.

To simplify the presentation in the numerical results to follow,  $\tilde{p}$  and  $p_{0z}$  are  
neglected (this has an impact on the physical solution, but not the mathematics  
as these terms are anti-symmetric). The initial condition is  $\zeta(x, y, z, 0) = 0$  and

the heating function is

$$H = H_1 H_2, \tag{4.1}$$

$$H_1 = .01 \sin^4(\pi y/L_1) \sin^2(\pi z/z_T), \tag{4.2}$$

$$H_2 = t_1 \sin(2\pi s_1/L_1) H_1, \tag{4.3}$$

212 where  $L_1 = 6000$  km is the size of the square horizontal domain,  $z_T = 12$  km  
 213 is the height at the top of the channel, the time factor  $t_1 = 1 - \exp(-t/86400)$ ,  
 214 the shift factor  $s_1 = x - u_0 t$  and  $u_0 = 10$  m/s. Note that the heating is  $O(1)$   
 215 in scaled terms, i.e., the magnitude is equal to the scaling value of  $W$ , the  $x$   
 216 derivative is of size  $W2\pi/L_1 \approx L^{-1}W$  and the storm is essentially the height  
 217 of the entire atmosphere. This heating consists of large scale warm and cold  
 218 air masses moving eastward at a velocity of 10 m/s. Note that the heating is  
 219 0 at the bottom and top of the channel, starts out slowly until essentially a  
 220 maximum is reached at 2 days. The results that follow will be shown at 4 days  
 221 (two days after maximum heating has been achieved). The grid sizes for models  
 222 are  $\Delta x = \Delta y = 100$  km,  $\Delta z = 1$  km,  $\Delta t = 40$  sec for the multiscale model and  
 223  $\Delta t = 1800$  sec for the reduced model.

224 Figure 1 shows the pressure perturbation from the multiscale model as a  
 225 function of time at the three horizontal grid points shown at the top of the plot  
 226 at a height of 3 km. Although the multiscale system has both low and high  
 227 frequencies present, they are clearly not activated as expected if the  $L_2$  norms  
 228 of the space and time derivatives of the nondimensional heating term are on the  
 229 order of the advective terms.

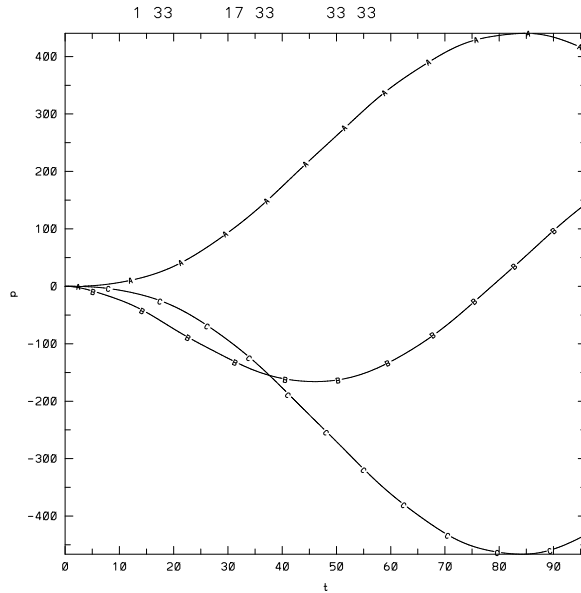


Figure 1: Multiscale pressure perturbation as a function of time at the  $(x,y)$  points at the top of the plot at  $z = 3$  km

230 Note that the multiscale and reduced systems are derived in completely  
 231 different manners. The multiscale hyperbolic system was mathematically proved  
 232 to describe the slowly evolving in time large scale atmospheric motions by a  
 233 continuum modification of system (2.1). The reduced system is derived from the  
 234 initialization constraints for (2.1) with the addition of a time dependent equation  
 235 for the vorticity. Because both independently are expected to describe the same  
 236 large scale slowly evolving solution a comparison of the solutions from models  
 237 based on the two different systems is of interest. Figure 2 on the following page  
 238 compares the results from the numerical model based on the multiscale system  
 239 (left hand side) and the numerical model based on the reduced system (right  
 240 hand side) for the variables shown at  $z = 9$  km and  $t = 4$  days. As expected  
 241 from the BDT, the solutions from the two models are quite similar. Although  
 242 only one level is shown, the relative  $l_2$  errors are 9.1%, 8.9% and 8.2% for the  
 243 horizontal divergence  $\delta$ , the vertical component of velocity  $w$  and the vertical

244 component of vorticity  $\zeta$ , respectively. As a number of terms of order  $\epsilon$  have  
245 been neglected, e.g., the term  $\delta\zeta$  in the vorticity equation in the derivation  
246 of the equation for  $w$ , these errors are completely reasonable. The lower level  
247 horizontal velocities in a primitive equation model grow unrealistically large in  
248 a few days and require an ad hoc boundary layer drag/dissipation to artificially  
249 slow down that growth (Sylvie Gravel, personal communication). Note that  
250 neither the multiscale or reduced model include any dissipation.

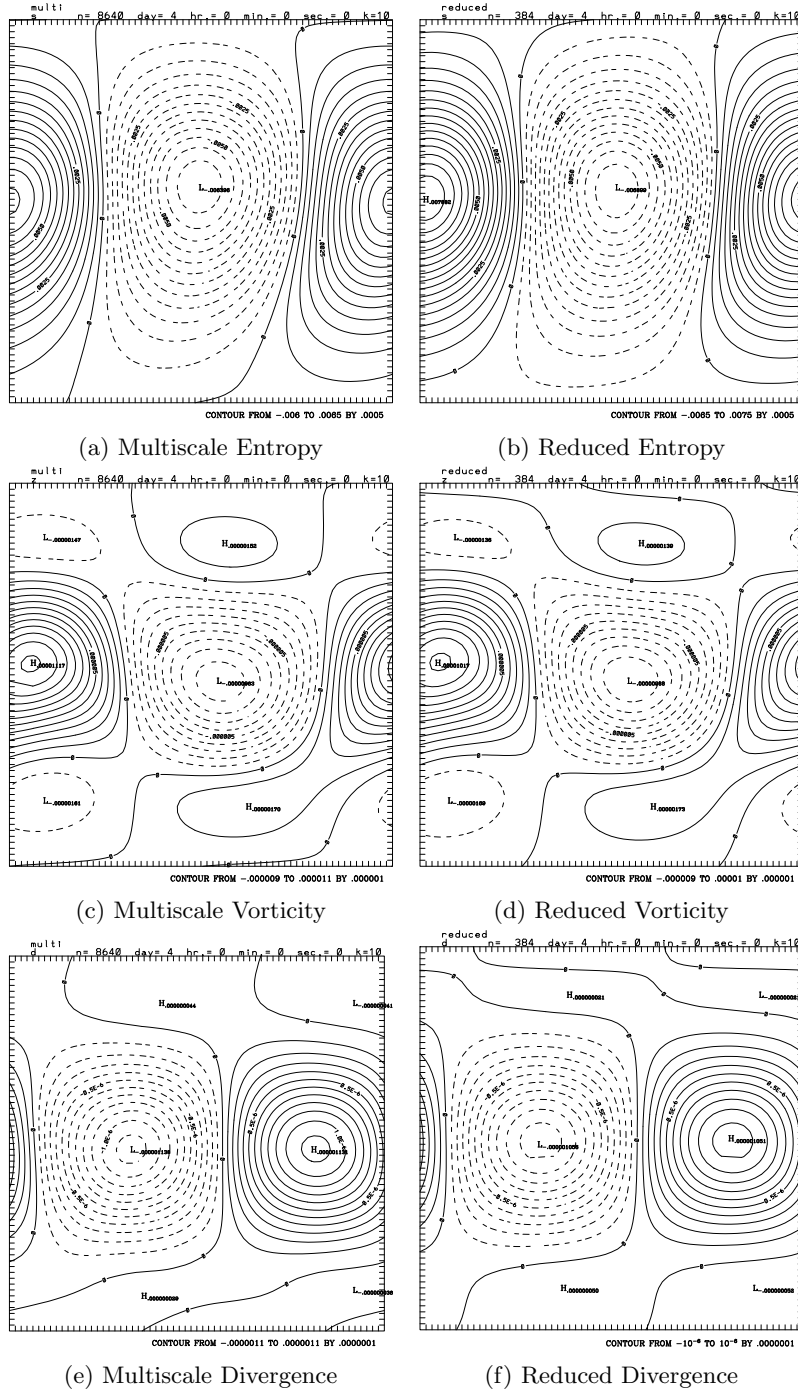


Figure 2: Comparison of multiscale model variables (left column) and reduced system model variables (right hand column) at  $z = 9$  km and  $t = 4$  days

251 There is a very important property of the elliptic equation for  $w$  that will be  
252 demonstrated next. Small scale noise at the lower boundary of that equation,  
253 e.g., noise caused by individual trees or rocks or small scale heating/cooling  
254 features that cause small changes in the vertical velocity at the surface, is not  
255 propagated very far into the solution. A solution of the elliptic equation for  $w$   
256 without any forcing term but with random noise at the surface was computed  
257 to show this property. Fig. 3 shows the random values of the vertical velocity  
258 at the surface. Fig. 4 and Fig. 5 show the resulting vertical velocity at the first  
259 and second levels of the model. Already at the second level the perturbations at  
260 the surface have been reduced by a factor of 10, i.e., the small scale irregularities  
261 at the lower boundary are not propagated very far into upper levels, exactly as  
262 expected from mathematical theory. This result should be contrasted with the  
263 sensitivity of the primitive equations at the surface that requires the ad hoc  
264 boundary layer treatment to prevent the rapid growth of perturbations at the  
265 surface.

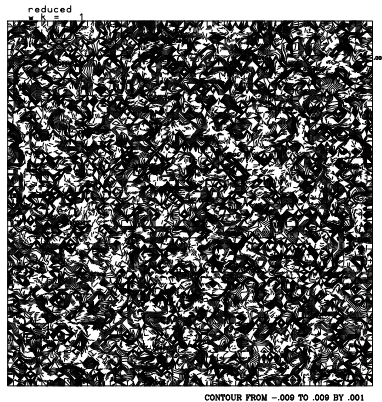


Figure 3: Random numbers for  $w$  at surface

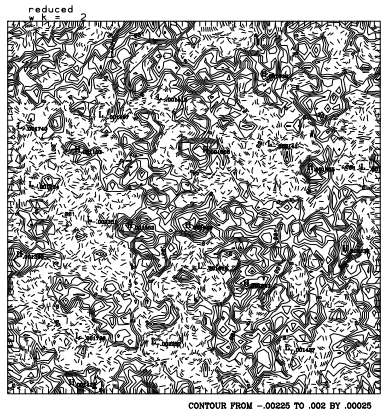


Figure 4:  $w$  at 1 km

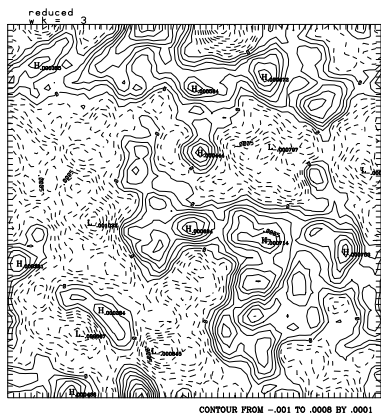


Figure 5:  $w$  at 2 km

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